

RADAR Titan Flyby during S91/T113

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- Sequence: s91
- Rev: 222
- Observation Id: t113
- Target Body: Titan
- Data Take Number: 271
- PDT Config File: S91_sip_port3_150717_pdt.cfg
- SMT File: smt_20150713.rpt
- PEF File: z0910b.pef

1 Introduction

This memo describes the Cassini RADAR activities for the T113 Titan flyby. This SAR data collection occurs during the S91 sequence of the Saturn Tour. This is a partial radar pass with inbound SAR imaging over the Xanadu area, and a ride-along outbound imaging segment. A sequence design memo provides the science context of the scheduled observations, an overview of the pointing design, and guidelines for preparing the RADAR IEB.

2 CIMS and Division Summary

CIMS ID	Start	End	Duration	Comments
222TI.T113WRMUP001_RIDER	2015-271T12:37:12	2015-271T21:03:12	08:26:00	
222TI.T113INALT001_PRIME	2015-271T21:03:12	2015-271T21:19:12	00:16:00	
222TI.T113INSAR001_PRIME	2015-271T21:19:12	2015-271T21:37:12	00:18:00	
222TI.T113RASAR001_PRIME	2015-271T21:37:12	2015-271T21:52:12	00:15:00	
222TI.T113RASAR002_RIDER	2015-271T21:37:12	2015-271T21:43:12	00:06:00	
222TI.T113OTALT001_PRIME	2015-271T21:52:12	2015-271T22:08:12	00:16:00	
222TI.T113OHSAR001_PRIME	2015-271T22:30:12	2015-271T23:14:12	00:44:00	

Table 1: t113 CIMS Request Sequence

Each RADAR observation is represented to the project by a set of requests in the Cassini Information Management System (CIMS). The CIMS database contains requests for pointing control, time, and data volume. The CIMS requests show a high-level view of the sequence design. Table 1 shows the CIMS request summary for this observation.

Division	Name	Start	Duration	Data Vol	Comments
a	Warmup	-9:00:0.0	08:15:0.0	29.5	Warmup
b	standard_radiometer_inbound	-0:45:0.0	00:05:0.0	0.3	radiometer quick-steps
c	scatterometer_compressed	-0:40:0.0	00:12:30.0	2.2	Compressed Scatt
d	standard_altimeter_inbound	-0:27:30.0	00:07:50.0	15.5	Inbound altimetry
e	standard_scatterometer_inbound	-0:19:40.0	00:00:4.0	0.6	Atmospheric Probe with Chirp
f	standard_scatterometer_inbound	-0:19:36.0	00:00:2.0	0.3	Atmospheric Probe with Tone
g	standard_sar_hi	-0:19:34.0	00:00:58.0	2.9	SAR Turn transition transition from scat, beam 3 only
h	standard_sar_pingpong	-0:18:36.0	00:02:36.0	34.3	Inbound SAR ping-pong
i	standard_sar_hi	-0:16:0.0	00:09:0.0	118.8	SAR-Hi Main Swath
j	standard_sar_hi	-0:07:0.0	00:09:0.0	118.8	SAR-Hi Main Swath
k	standard_sar_hi	00:02:0.0	00:05:0.0	66.0	SAR-Hi Ride-Along
l	scatterometer_compressed	00:07:0.0	00:12:0.0	2.2	Compressed Scatt (off target)
m	standard_sar_hi	00:19:0.0	00:01:0.0	3.0	SAR Turn transition transition from scat, beam 3 only
n	standard_sar_hi	00:20:0.0	00:02:0.0	27.6	Outbound bonus SAR
o	standard_sar_hi	00:22:0.0	00:00:52.0	2.6	SAR Turn transition transition from SAR, beam 3 only
p	standard_scatterometer_outbound	00:22:52.0	00:00:4.0	0.6	Atmospheric Probe with Tone
q	standard_scatterometer_outbound	00:22:56.0	00:00:2.0	0.3	Atmospheric Probe with Chirp
r	standard_altimeter_outbound	00:22:58.0	00:05:50.0	11.6	Outbound altimetry
s	standard_sar_hi	00:28:48.0	00:01:24.0	4.2	SAR Turn transition transition from alt, beam 3 only
t	scatterometer_compressed	00:30:12.0	00:09:48.0	1.8	Compressed Scatt (off target)
u	scatterometer_imaging	00:40:0.0	00:08:0.0	30.7	Outbound scatterometer imaging
v	scatterometer_imaging	00:48:0.0	00:17:0.0	65.3	Outbound scatterometer imaging
w	scatterometer_imaging	01:05:0.0	00:10:0.0	38.4	Outbound scatterometer imaging
x	standard_radiometer_outbound	01:15:0.0	00:25:0.0	1.5	Outbound radiometry
Total				578.8	

Table 2: Division summary. Data volumes (Mbits) are estimated from maximum data rate and division duration.

Div	Alt (km)	Slant range (km)	B3 Size (target dia)	B3 Dop. Spread (Hz)
a	174268	off target	0.22	off target
b	12822	off target	0.02	off target
c	11227	off target	0.02	off target
d	7302	7303	0.01	902
e	4942	4942	0.01	1195
f	4923	4923	0.01	1198
g	4913	4913	0.01	1200
h	4633	4715	0.01	1248
i	3898	3943	0.01	1392
j	1738	1777	0.01	2100
k	1098	1255	0.00	2469
l	1738	2531	0.01	2101
m	4749	off target	0.01	off target
n	5040	5766	0.01	1179
o	5631	5805	0.01	1092
p	5891	5891	0.01	1058
q	5911	5911	0.01	1055
r	5921	5921	0.01	1054
s	7704	7704	0.01	865
t	8139	8994	0.01	828
u	11228	13341	0.02	632
v	13784	15136	0.02	524
w	19262	20020	0.03	373
x	22499	23063	0.03	313

Table 3: Division geometry summary. Values are computed at the start of each division. B3 Doppler spread is for two-way 3-dB pattern. B3 size is the one-way 3-dB beamwidth

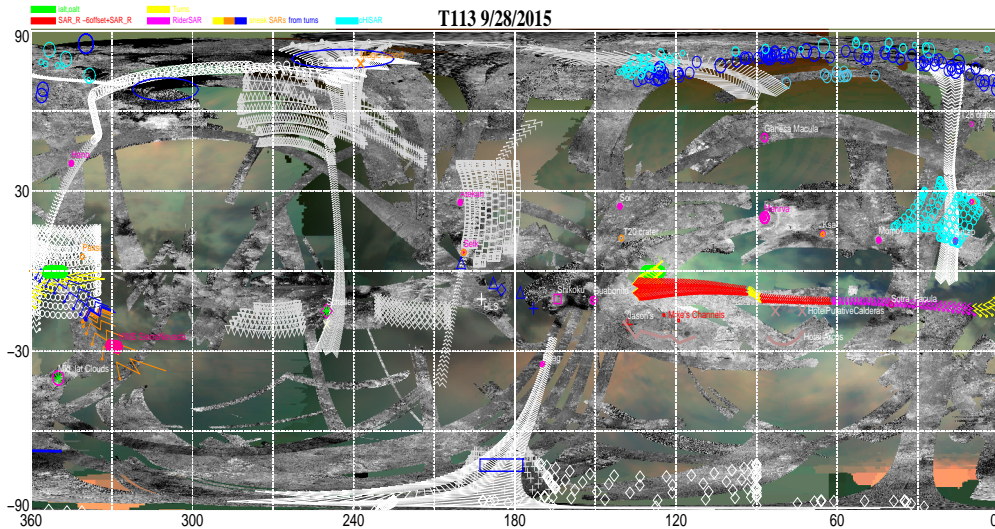


Figure 1: Coverage areas overlaid on Titan map showing prior optical and radar imaging.

Although the CIMS requests show Low-SAR intervals, in reality the radar will be operated in Hi-SAR mode through most of this flyby.

The CIMS requests form the basis of a pointing design built using the project pointing design tool (PDT). The details of the pointing design are shown by the PDT plots on the corresponding tour sequence web page. (See <https://cassini.jpl.nasa.gov/radar>.) The RADAR pointing sequence is ultimately combined with pointing sequences from other instruments to make a large merged c-kernel. C-kernels are files containing spacecraft attitude data.

A RADAR tool called RADAR Mapping and Sequencing Software (RMSS) reads the merged c-kernel along with other navigation data files, and uses these data to produce a set of instructions for the RADAR observation. The RADAR instructions are called an Instrument Execution Block (IEB). The IEB is produced by running RMSS with a radar config file that controls the process of generating IEB instructions for different segments of time. These segments of time are called divisions with a particular behavior defined by a set of division keywords in the config file. Table 2 shows a summary of the divisions used in this observation. Table 3 shows a summary of some key geometry values for each division.

3 Overview

T113 is a partial pass. The observation starts with inbound altimetry and an atmospheric probe before turning to the inbound SAR imaging of the Xanadu area. After closest approach, INMS becomes prime and radar obtains and additional 5 minutes of SAR ride-along imaging. Then radar becomes prime again and we turn back on target for an outbound altimeter segment and atmospheric probe. Following this is a high altitude imaging segment with 8 scan lines providing SAR imaging that fills in some unobserved areas North of Sinlap.

4 Mode Specific Operation and Performance

Many details of standard radar sequencing during the 4 main modes (Radiometry, Scatterometry, Altimetry, and SAR) have been discussed in previous sequence memos for prior observations. Refer to these for details. Some selected performance highlights are illustrated in figures and explained in the following subsections.

4.1 Coverage Layout

Figure 1 shows the layout of the different T113 data collections on a map of Titan. The red jagged lines show the beam centers of the ride-along SAR swath as it sweeps across the Xanadu area. The cyan symbols show the high altitude

imaging North of Sinlap. The green symbols show the altimeter tracks.

4.2 SAR Resolution Performance

For all of the SAR divisions the effective resolution can be calculated from the same equations used in the high-altitude imaging discussion. Figure 2 shows the results from these equations using the parameters from the IEB as generated by RMSS. The calculations are performed for the boresight of beam 3 which is the center of the swath.

Projected range increases with decreasing incidence angle, so the range resolution varies across the swath with better resolution at the outer edge. The SAR pointing profile decreases the incidence angle as time progresses and altitude increases, so there is progressive deterioration of range resolution away from closest approach. The projected range resolution rapidly deteriorates as the incidence angle decreases toward zero at the very beginning and end of the swath and during the close approach altimetry segment.

Azimuth resolution is a function of the synthetic aperture size which is determined by the length of the receive window in each burst (assuming the receive window is always filled with echos). Azimuth resolution deteriorates less quickly because the number of pulses and the length of the receive window are increased as altitude increases which mitigates the increasing doppler bandwidth of the beam patterns. The receive window length increases to fill the round trip time until the science data buffer is filled. At this point it is no longer possible to extend the receive window, and azimuth resolution starts to deteriorate more rapidly.

5 Revision History

1. Jun 17, 2016: Final release

6 Acronym List

ALT	Altimeter - one of the radar operating modes
BAQ	Block Adaptive Quantizer
CIMS	Cassini Information Management System - a database of observations
Ckernel	NAIF kernel file containing attitude data
DLAP	Desired Look Angle Profile - spacecraft pointing profile designed for optimal SAR performance
ESS	Energy Storage System - capacitor bank used by RADAR to store transmit energy
IEB	Instrument Execution Block - instructions for the instrument
ISS	Imaging Science Subsystem
IVD	Inertial Vector Description - attitude vector data
IVP	Inertial Vector Propagator - spacecraft software, part of attitude control system
INMS	Inertial Neutral Mass Spectrometer - one of the instruments
NAIF	Navigation and Ancillary Information Facility
ORS	Optical Remote Sensing instruments
PDT	Pointing Design Tool
PRI	Pulse Repetition Interval
PRF	Pulse Repetition Frequency
RMSS	Radar Mapping Sequencing Software - produces radar IEB's
SAR	Synthetic Aperture Radar - radar imaging mode
SNR	Signal to Noise Ratio
SOP	Science Operations Plan - detailed sequence design
SOPUD	Science Operations Plan Update - phase of sequencing when SOP is updated prior to actual sequencing
SSG	SubSequence Generation - spacecraft/instrument commands are produced
SPICE	Spacecraft, Instrument, C-kernel handling software - supplied by NAIF to use NAIF kernel files.
TRO	Transmit Receive Offset - round trip delay time in units of PRI

SAR Surface Resolution

epoch = 2015-09-28T21:37:11.818

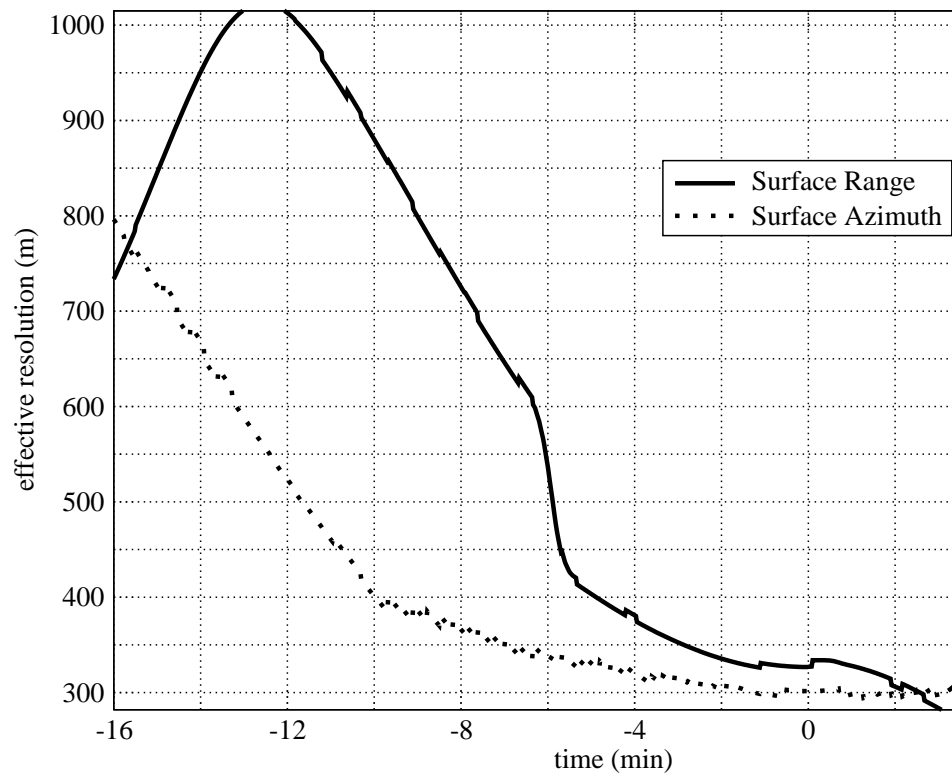


Figure 2: SAR projected range and azimuth resolution. These values are computed from the IEB parameters and are not related to the pixel size in the BIDR file. The pixel size was selected to be always smaller than the real resolution.